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Electron and nuclear spin dynamics in the thermal mixing model of dynamic nuclear polarization

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A novel mathematical treatment is proposed for computing the time evolution of dynamic nuclear polarization processes in the low temperature thermal mixing regime. Without assuming any *a priori* analytical form for the electron polarization, our approach provides a quantitative picture of the steady state that agrees with the well known Borghini prediction based on thermodynamic arguments, as long as the electrons–nuclei transition rates are fast compared to the other relevant time scales. Substantially different final polarization levels are achieved instead when the latter assumption is relaxed in the presence of a nuclear leakage term, even though very weak, suggesting a possible explanation for the deviation between the measured steady state polarization and relaxation times, once the electrons/nuclei concentration ratio and the typical rates of the microscopic processes involving the two spin species are specified. Numerical results are shown to account for the manifold dynamic behaviours of typical DNP samples.

I. Introduction

Dynamic Nuclear Polarization (DNP) techniques are attracting increasing interest due to their proven ability to improve the sensitivity of Nuclear Magnetic Resonance (NMR) experiments by several orders of magnitude, not only in the solid state¹ but also in solution.² Although the disruptive potential of DNP is nowadays well known and accepted, the articulated physical scenario underlying these complex phenomena is still puzzling the scientific community.

In the solid state, three different DNP regimes can be specified according to the resonance frequency ω_n of the nuclei to be polarized, the width $\Delta \omega_e$ of the Electron Spin Resonance (ESR) line of the free radicals used as polarization source and the magnitude of dipolar coupling between those radicals $1/T_{2e}$.

When $\omega_n > \Delta \omega_e$ the main mechanism for nuclei to polarize, known as the Solid Effect (SE), proceeds *via* microwave (MW) assisted forbidden transitions involving simultaneous flip-flops or flip-flips of one electron and one nucleus in mutual dipolar interaction. The SE model has been extensively described in ref. 3 and the role played by spin diffusion was discussed in ref. 4. More recently, an exact quantum mechanical treatment of the SE based on the density matrix approach has been proposed by Hovav *et al.*^{5,6} and by Karabanov and collaborators.⁷

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On the other side, when $\omega_n < \Delta \omega_e$, nuclei can flip between different Zeeman levels within an energy conserving three particle mechanism involving a simultaneous flip-flop of two electron spins (referred to as the ISS process hereafter). Such a process, driven by electron-electron and electron-nucleus time dependent dipolar interactions, does not involve forbidden transitions and it is thus in general more effective than the SE. As far as the typical interaction time T_{2e} between different paramagnetic centres is long compared to the electron spin lattice relaxation time T_{1e} (*i.e.* the unpaired electrons are – on average – relatively far from each other), the polarization mechanism is referred to as the Cross Effect (CE) and was first analyzed in ref. 8-11 The CE model has been successfully exploited to describe DNP with bi-radicals,^{12,13} i.e. polarization procedures where the polarizing agents are tailored molecules carrying two unpaired electrons having resonance frequencies differing exactly by $\omega_{\rm p}$. Similarly to the SE, also the CE model has been rigorously computed by ab initio quantum mechanical techniques.14,15

However, the novel applications, which in the last decade have renewed the attention on DNP due to their potential impact on biosciences,^{16,17} are based on samples which generally fall into none of the two models previously described. The nuclei of interest for this kind of applications (typically ¹³C or ¹⁵N) have a low-gyromagnetic ratio γ_n and, as a consequence, a resonance frequency ω_n significantly smaller than the width of the typical radicals used as polarizing agents (trityls or nitroxides). Moreover, the concentration of these radicals, usually above 10 mM in the solution to polarize, leads to non-negligible electron spin–spin interactions, *i.e.* to the condition $1/T_{2e} \gg 1/T_{1e}$ that defines the Thermal Mixing (TM) regime. Under TM assumptions,

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when a transition is saturated by an external, frequency selective, microwave field, the whole electron spin distribution reacts through energy conserving flip-flops (spectral diffusion) and evolves towards a new steady state different from Boltzman equilibrium. Hence, via the previously introduced ISS processes, the nuclear Zeeman populations are also perturbed and. depending on the frequency of the saturated electron transition, possibly result in depletion or enhancement of the population of the ground state and thus in an enhanced nuclear magnetic order. At high temperature, where a linear expansion of the density matrix can be used, the evolution of the electron and nuclear polarizations is accounted for by a set of rate equations formulated by Provotorov.^{3,18–21} At the typical temperatures (about 1 K) where most of the DNP experiments aimed to obtain substantial nuclear polarization values (>0.1) are performed, however, a linear expansion of the density matrix is not allowed and consequently the Provotorov approach does not apply. Fortunately under these conditions, the electron resonance lines of the free radicals used as polarizing agents are normally inhomogeneously broadened and can likely be depicted as the convolution of several individual packets of given resonance frequency, mutually connected by the electronelectron dipolar interaction.³ Hyperpolarization via TM at low temperature has been first discussed by Borghini, who also calculated the steady state solution for the electron and nuclear polarization after imposing certain constraints to the model.²² The same - steady state - result has been achieved by Abragam and Goldman³ using a slightly different mathematical procedure. The Borghini prediction has the merit of reproducing, at least qualitatively, many observations obtained in DNP experiments at low temperature.²³⁻²⁵ It has, at the same time, several limitations, the major of them are listed below.

(i) It provides a picture of the steady state polarization without contributing in any way to the understanding of the time evolution of nuclear spin order towards equilibrium.

(ii) It is obtained by assuming an *a priori* analytical form for the electron spin polarization.

(iii) It is derived under strong saturation conditions and in the limit of perfect contact between the electron and the nuclear reservoirs; the latter hypothesis consists in assuming highly effective *ISS* processes.

As a consequence of the constraints (ii) and (iii), the model leads to a final steady state nuclear polarization which substantially depends only on the lattice temperature, the magnetic field in which the DNP phenomenon takes place and the width of the electron resonance line. The sample specific parameters like electron and nuclear spin-lattice relaxation rates and the relative concentration of the two spin species play only a secondary role.

In this work we aim to overcome these three limitations by introducing a dynamic analysis of the low temperature TM model (described in detail in Section II) based on a set of rate equations presented in Section III which spontaneously provide the time evolution laws of the electron and nuclear polarizations without any *a priori* assumption of their functional form. The results obtained by numerically solving the rate equations are reported in Section IV and discussed in Section V, with particular emphasis on the modifications occurring in the steady state and in the dynamic parameters in the presence of nuclear leakage and finite electron–nucleus exchange. The technical arguments underlying the derivation of the rate equations set are given in Appendices A and B, while in Appendix C, for the convenience of the reader, the derivation of the Borghini equation for the steady state as proposed by Abragam and Goldman³ is briefly recalled.

II. Model description

A system made up of N_n nuclear spins I with Larmor frequency ω_n and N_e electron spins S with mean Larmor frequency ω_e is considered (S = I = 1/2). Typically ω_e is about three orders of magnitude higher than ω_n . Since all terms in the nuclear Hamiltonian with the exception of the Zeeman one are small in comparison to the other energy scales considered, nuclei are assumed to resonate all at the same frequency.

The first assumption of the model is that the main contribution to the ESR line shape is the spread of *g*-factors (inhomogeneous broadening of the I type), which reflects the different orientation of the unpaired magnetic moments with respect to the external magnetic field due to single ion anisotropy. The practical relevance of this assumption is confirmed, for instance, by the experimental evidence obtained from the widely used trityl radical family.²³ Inhomogeneously broadened lines are conveniently decomposed in a sequence of narrow individual spin packets of frequency $\omega_i = \omega_e - \Delta_i$, width $\delta\omega$ and relative weight f_i (see Fig. 1) such that:

$$\sum_{i} f_{i} = 1 \tag{1}$$
$$\sum_{i} f_{i} \Delta_{i} = 0,$$

valid for any ESR line. The continuous limit is recovered when $\delta \omega \rightarrow 0$. For each electron packet, a local polarization $P_{\text{e},i} = 2\langle S_z^i \rangle$ may be defined, where $\langle \rangle$ stands both for the quantum mechanic expectation value and the average over all electron



Fig. 1 Discretization of the ESR line. Each basic packet centered at $\omega_i = \omega_e - \Delta_i$ is characterized by a width $\delta\omega$ and a weight f_i . By way of example the irradiation frequency is set to $\omega_0 = \omega_e - \Delta_0$ and the nuclear Larmor frequency corresponds to three basic packets ($\delta n_p = 3$).

spins belonging to the *i*-th packet. The second assumption of the model is that ω_n is smaller than the ESR line width, as actually is the case for low γ_n nuclei with almost any type of radical or for high gamma nuclei, such as ¹H, with broad line radicals (*e.g.* nitroxides). In view of the mathematical description proposed in the next section, it is useful to introduce at this point the variable

$$\delta n_{\rm p} = \frac{\omega_{\rm n}}{\delta \omega},\tag{2}$$

i.e. the number of electron packets corresponding to the nuclear Larmor frequency (see Fig. 1).

The third assumption is that the spin dynamics of the system is governed only by the five processes depicted in Fig. 2 and briefly explained here below.

Microwave irradiation

Single electron transitions are stimulated by an external MW field at $\omega_0 = \omega_e - \Delta_0$ that leads the electron spin system out of equilibrium (panel A of Fig. 2). The characteristic time of this



Fig. 2 Microscopic events in the thermal mixing model. Electron spins belonging to a generic packet *i* are represented in black with a subscript e,*i*, nuclear spins in white with a subscript n. Panel A: electron flip due to MW irradiation; panel B: spectral diffusion (δ is a generic number of packets); panel C: *ISS* processes; panel D: electron spin–lattice relaxation; panel E: nuclear spin–lattice relaxation ("leakage"), under the assumption $P_{0,n} \approx 0$ (see eqn (4)).

process is named T_{1MW} . For high microwave power $T_{1MW} \rightarrow 0$ and the packet Δ_0 is saturated.

Spectral diffusion

The basic transition of this process, which involves only electrons and conserves both the total energy and the total electron polarization, is represented in Fig. 2, panel B. Spectral diffusion transitions are promoted by the dipolar interaction among electrons and are characterized by a time constant T_{2e} which in the thermal mixing regime is assumed to be much shorter than any other relevant time scale (in practice, in a solid solution used for DNP, it is typically <1 μ s²³).

ISS process

The mechanism is sketched in Fig. 2, panel C, and consists in a simultaneous flip-flop of two electron spins, belonging to packets separated by ω_n , compensated by a nuclear spin flip. This process allows an energy transfer between nuclei and electrons and while the total energy and the electron polarization are conserved, the nuclear polarization is not. In a real system the ISS process, being promoted by dipolar interactions, involves only nuclei which are sufficiently close to paramagnetic centres. Thanks to nuclear spin diffusion, the local information is then spread throughout the entire system. A model which takes into account both these aspects (local dipolar interaction and spatial diffusion) would be very cumbersome to treat, because of the $N_{\rm p}$ degrees of freedom. In the present work we adopt a mean field approach where all possible terms comprising a nucleus and two electrons flip with a characteristic time $T_{\rm ISS}^{\rm eff}$ which does not depend on the mutual distances between the three particles. Setting low values of $T_{\rm ISS}$ the model mimics the fast spin diffusion limit, while setting high values of $T_{\rm ISS}$ the model mimics the slow spin diffusion limit.

Electron spin-lattice relaxation

The process, shown in Fig. 2, panel D, accounts for the contact between the electron system and the lattice which leads $P_{e,i}(t)$ towards the electron thermal equilibrium polarization P_0 :

$$P_0 = -\tanh\frac{\hbar\omega_{\rm e}}{2k_{\rm B}T},\tag{3}$$

where T is the temperature of the lattice. Its characteristic time constant T_{1e} ranges from hundreds of milliseconds to a few seconds,^{23,26} for $T \leq 10$ K.

Nuclear spin-lattice relaxation ("leakage")

The process is shown in Fig. 2, panel E. The nuclear system is directly in contact with the lattice *via* slow processes (where electrons are not involved), with a characteristic time $T_{\rm ln} > 10^3$ s, also called "leakage" terms. Analogously to the latter case, these processes lead nuclear polarization, $P_{\rm n}(t)$, towards the thermal equilibrium value $P_{0,\rm n}$:

$$P_{0,n} = \tanh \frac{\hbar \omega_n}{2k_{\rm B}T} \approx 0. \tag{4}$$

In the next section, the temporal evolution of the nuclear polarization for the model described above will be determined by a closed set of equations involving only $P_n(t)$ and $P_{e,i}(t)$.

III. Rate equations approach

To investigate the dynamic evolution of the model summarized in the previous section, we introduce a system of rate equations and solve them numerically. The discrete time step dt for the numerical integration is defined as the inverse of the sum of the single rates for the five processes sketched in Fig. 2.

Microwave rate

The rate of all possible microwave events is $W_{\rm MW} = N_{\rm e} f_0 / T_{\rm 1MW}$.

ISS process rate

In the mean field approximation, the number of the possible processes involving two electrons and one nucleus is given by $\sum f_i N_e f_{i+\delta n_p} N_e N_n$. The total *ISS* rate needs to linearly scale with the size of the system in order to assure a correct thermodynamic limit. To achieve this, as is usually done for fully connected models, the effective time constant of each three particle process must depend on the system size and scale as:

$$T_{\rm ISS}^{\rm eff} = T_{\rm ISS} N_{\rm e} N_{\rm n}, \tag{5}$$

where the constant T_{ISS} is size independent. Then the total rate of all *ISS* events is $W_{\text{ISS}} = N_{\text{e}} \sum f_i f_{i+\delta n_p} / T_{\text{ISS}}$.

Electron and nuclear spin-lattice relaxation rate

The rate of all the electron spin-lattice relaxation processes is $W_{\rm e} = N_{\rm e}/T_{\rm 1e}$. Similarly the rate of all the nuclear spin-lattice relaxation processes is $W_{\rm n} = N_{\rm n}/T_{\rm 1n}$.

Spectral diffusion rate

Being proportional to $1/T_{2e}$, it is assumed to be by far the highest total transition rate among those characterizing the five considered mechanisms. A large rate makes the time step dt very small and consequently produces a dramatic slow-down of the simulation procedure. To avoid that, the dynamic problem has been separated into two steps: a "Short-Term Thermalization" (STT) step involving spectral diffusion and a "Long-Term Evolution" (LTE) where the system evolves according to defined differential equations under the action of the remaining processes. Depending on the value of T_{1MW} , the microwave pumping can be considered either as a fast process (contributing to STT) or as a player in the LTE. In the first case the time step is defined as:

$$dt = \frac{1}{W_0 + W_{ISS} + W_n},\tag{6}$$

while in the second case one has:

$$dt = \frac{1}{W_{\rm e} + W_{\rm ISS} + W_{\rm n} + W_{\rm 1MW}}.$$
 (7)

The numerical procedure can be summarized as follows:

$$\begin{cases} P_{\mathrm{e},i}^{0}(t) \stackrel{\mathrm{STT}}{\Rightarrow} \begin{cases} P_{\mathrm{e},i}(t) \stackrel{\mathrm{LTE}}{\Rightarrow} \\ P_{\mathrm{n}}(t) \end{cases} \begin{cases} P_{\mathrm{e},i}^{0}(t+\mathrm{d}t) \\ P_{\mathrm{n}}(t+\mathrm{d}t) \end{cases}$$

A. STT. The goal of this step is to "thermalize" the profile of the electron polarization $P_{e,i}^{0}(t)$ at a generic time *t*, under the action of those processes considered "fast". The steady state profile $P_{e,i}(t)$ can be obtained by imposing the detailed balance condition (see Appendix A):

$$P_{e,i}(t) = -\tanh[\beta(\Delta_i - c)].$$
(8)

This equation depends on two parameters that can be computed using conservation principles. In this respect we need to discuss separately the two cases where spectral diffusion only or both spectral diffusion and microwave promoted processes are viewed as fast.

1. Spectral diffusion only. This case is relevant for relaxation experiments (with microwave field off by definition) or when a non-saturating microwave field is applied in a DNP experiment. To calculate the two parameters c and β , the conservation of *both* the total energy and the total polarization is imposed:

$$\sum f_i \Delta_i (P_{e,i}(t) - P_{e,i}^0(t)) = 0$$
(9)

$$\sum f_i (P_{e,i}(t) - P_{e,i}^0(t)) = 0.$$
 (10)

2. Spectral diffusion and saturating microwaves. For high irradiation power, the microwaves act as an infinite bath for energy exchange and force the polarization $P_{e,0}(t)$ to 0 and in turn, by eqn (8), $c = \Delta_0$. In order to evaluate the second parameter β , the output of the STT step is conveniently written as:

$$P_{e,i}(t) = P_{e,i}^0(t) + \delta P_{e,i} \text{ if } i \neq 0$$
 (11)

$$P_{e,0}(t) = P_{e,0}^{0}(t) + \delta P_{e,0} + \delta P^{MW}$$
(12)

where $\delta P_{e,i}$ and δP^{MW} are the variations of the polarization of the *i*-th packet induced by spectral diffusion and by micro-waves respectively. The conservation of the total energy and of the total polarization can now be written as:

$$\sum f_i \Delta_i \delta P_{\mathrm{e},i} = 0 \tag{13}$$

$$\sum f_i \delta P_{\mathbf{e},i} = 0. \tag{14}$$

Eqn (14) can be recast as $f_0 \delta P_{e,0} = -\sum_{i \neq 0} f_i \delta P_{e,i}$, so that:

$$\sum f_i \Delta_i \delta P_{\mathbf{e},i} = \sum_{i \neq 0} f_i \Delta_i \delta P_{\mathbf{e},i} + f_0 \Delta_0 \delta P_{\mathbf{e},0}$$

$$= \sum f_i (\Delta_i - \Delta_0) P_{\mathbf{e},i} = 0.$$
(15)

Summing on both sides $\sum f_i(\Delta_i - \Delta_0)P_{e,i}^0(t)$ we get:

$$\sum f_i (\Delta_i - \Delta_0) P_{e,i}(t) = \sum f_i (\Delta_i - \Delta_0) P_{e,i}^0(t).$$
(16)

The condition for β

$$\sum f_i \{ P_{e,i}^0(t) + \tanh[\beta(\Delta_i - \Delta_0)] \} (\Delta_i - \Delta_0) = 0 \quad (17)$$

identifies the unique solution for $P_{e,i}(t)$.

B. LTE. In this second step the profile of the electron polarization $P_{e,i}^0(t + dt)$ at time t + dt is deduced from the output of the STT. When microwaves are off (relaxation) or they have been already taken into account in the STT, the rate

equations for $P_{e,i}$ and P_n , as discussed in Appendix B, can be set down in the form:

$$P_{e,i}^{0}(t+dt) = P_{e,i}(t) + dt \left(\frac{P_0 - P_{e,i}(t)}{T_{1e}} + \frac{f_{i-\delta n_p} \Pi_- + f_{i+\delta n_p} \Pi_+}{2T_{ISS}} \right)$$
(18)

$$P_{\mathrm{n}}(t+\mathrm{d}t) = P_{\mathrm{n}}(t) + \mathrm{d}t \left[\frac{P_{\mathrm{0n}} - P_{\mathrm{n}}(t)}{T_{\mathrm{1n}}} - \frac{N_{\mathrm{e}}}{2T_{\mathrm{ISS}}N_{\mathrm{n}}} \sum f_{i}f_{i+\delta n_{\mathrm{p}}}\Pi_{\mathrm{n}} \right]$$

where $\Pi_{-} = \Pi_{-}(i,t)$, $\Pi_{+} = \Pi_{+}(i,t)$, $\Pi_{n} = \Pi_{n}(i,t)$ are given by the expressions:

$$\Pi_{-} = P_{e,i-\delta n_{p}}(t) - P_{e,i}(t) + P_{n}(t)[1 - P_{e,i-\delta n_{p}}(t)P_{e,i}(t)]$$

$$\Pi_{+} = P_{e,i+\delta n_{p}}(t) - P_{e,i}(t) + P_{n}(t)[1 - P_{e,i+\delta n_{p}}(t)P_{e,i}(t)]$$

$$\Pi_{n} = P_{e,i+\delta n_{p}}(t) - P_{e,i}(t) + P_{n}(t)[1 - P_{e,i+\delta n_{p}}(t)P_{e,i}(t)]$$

The effect of non-saturating MW irradiation can be included by the term $-P_{e,0}(t)/T_{1MW}$ in the rate equation for $P_{e,0}(t)$, describing the electron polarization of the irradiated packet.

IV. Numerical results

In this section we exploit the formalism described above to calculate, in specific situations, those curves which are normally measured in DNP experiments, namely build-up and relaxation curves. The so-called build-up curve describes the polarization growth over time under selective microwave irradiation and it is characterized by two parameters, the final nuclear polarization P_n :

$$P_{\rm n} = P_{\rm n}(t \to \infty), \qquad (19)$$

and the characteristic time $T_{\rm pol}$ that we define by the equation:

$$P_{\rm n}(t=T_{\rm pol}) = P_{\rm n}\left(1-\frac{1}{e}\right),\tag{20}$$

with $P_n(t = 0) = 0$. T_{pol} corresponds to the usual time constant in the case of exponential build-up but, through eqn (20), one has a more general quantification of the growing speed regardless of the specific form of the polarization function.

Relaxation curves, on the other hand, represent the spontaneous (at microwaves off) equilibrium recovery of a system that is prepared out of equilibrium at t = 0. The characteristic time is named T_{relax} and, if one starts with a fully polarized system ($P_n(t = 0) = 1$), the constant T_{relax} is defined by:

$$P_{\rm n}(t = T_{\rm relax}) = P_{0,\rm n}\left(1 - \frac{1}{e}\right) + \frac{1}{e},$$
 (21)

where $P_{0,n}$ is defined in eqn (4).

Setting of the parameters

In order to make the results of our simulation relevant for a deeper understanding of experimental data, a number of basic parameters are borrowed from some of the best known DNP samples for biomedical applications. In this framework, the typical number of ¹³C nuclei (that in most cases are the object of DNP) is in the range 5–15 M, while the radical

concentration is 10–20 mM.²⁷ It is thus reasonable, for our purposes, to set $N_n/N_e = 1000$. Concerning the distribution of the electron spin resonances, it looks worth using a Gaussian function with a full width at half maximum $\Delta \omega_e = 63$ MHz, as surrogate of the ESR line of the trityl doped samples studied in ref. 23. This is actually one of the few ESR line widths which have been measured in a standard DNP environment (a magnetic field of 3.35 T and a temperature of 1.2 K) and that turned out to be inhomogeneously broadened by *g*-factor spreading. Finally, despite the fact that our approach is not restricted to the case of saturating microwaves, we work under the approximation $T_{1\text{MW}} \rightarrow \infty$ to better mimic what is normally done in actual experiments.

ESR line discretization

The Gaussian function used as a model of the ESR line is truncated at 3σ , where $\sigma = 27$ MHz. This defines a frequency interval of about 160 MHz to be covered by electron spin packets. Three different discretizations of the line are employed, with different packet widths: $\delta\omega = 32$, 10.6 and 6.4 MHz. The corresponding number of packets N_p is thus equal to 5, 15 and 25. The value of ω_n is set at 32 MHz (resonance frequency of ¹³C nuclei at the typical DNP magnetic field of 3.35 T) which corresponds, respectively, to $\delta n_p = 1,3$ and 5 for the three selected values of N_p . The MW frequency is set equal to $\omega_0 = \omega_e - \sigma$, which corresponds to saturating the packet 2 (for $N_p = 5$), 5 (for $N_p = 15$) and 8 (for $N_p = 25$). For this value of ω_0 the final nuclear polarization is maximal.

The results we show are obtained by integrating the rate equations system (eqn (18)) together with the STT step defined by the detailed balance condition (eqn (8)), with $c = \Delta_0$ and β set by eqn (17). The boundary conditions of the ESR line shape are correctly implemented by imposing $f_{i+\delta n_p} = 0$ ($f_{i-\delta n_p} = 0$) if $i + \delta n_p > N_p$ ($i - \delta n_p < 0$).

A. Fast *ISS* limit. The case of highly effective contact between electrons and nuclei, which corresponds to very fast *ISS* processes compared to electron spin lattice relaxation $(T_{\rm ISS} \ll T_{\rm 1e})$, is considered first, by setting $T_{\rm ISS} = 10^{-3}T_{\rm 1e}$ and $T_{\rm 1e} = 1$ s (the latter from ref. 23) in our computational tool.

In Fig. 3, three different build-up curves computed with an increasing value of $N_{\rm p}$ and in the absence of leakage $(T_{\rm 1n} \rightarrow \infty)$ are presented. The continuum limit is approached very fast and a good convergence is reached already with $N_{\rm p} = 5$. The final nuclear polarization is $P_{\rm n} = 0.825$ and the characteristic time constant of the polarization build-up curves $T_{\rm pol}$ is equal to 1140 s.

In Fig. 4, three relaxation curves, describing the evolution of the nuclear polarization in the absence of both microwaves and leakage and calculated with a different discretization of the ESR line, are shown. Again, the curve with $N_{\rm p} = 5$ is already representative of the continuous limit. The characteristic time $T_{\rm relax} \approx 15600$ s, evaluated by means of eqn (21) with $P_{0,\rm n} = 0$, turns out to be about 10 times longer than the corresponding $T_{\rm pol}$.

In Fig. 5 the effect of nuclear leakage is investigated. The thermal contact between the nuclear spin system and the lattice significantly affects the build-up curve even when T_{1n} is much slower than any other transition considered. For $T_{1n} = 10\,000$ s, the final nuclear polarization goes down to $P_n = 0.707$, which corresponds to a reduction of 15% with



Fig. 3 Polarization build-up curves at 3.35 T and 1.2 K of a DNP system characterized by the following parameters: $N_n/N_e = 1000$, $T_{1e} = 1$ s, $T_{ISS} = 0.001$ s, $T_{1n} \rightarrow \infty$, for different discretizations of the ESR line: $N_p = 5$ (solid line), $N_p = 15$ (dashed line), $N_p = 25$ (dotted line). The final nuclear polarization is $P_n = 0.825$ reached with a characteristic time constant $T_{pol} = 1140$ s.



Fig. 4 Relaxation curves at 3.35 T and 1.2 K of a DNP system with $N_n/N_e = 1000$, $T_{1e} = 1$ s, $T_{ISS} = 0.001$ s, $T_{1n} \rightarrow \infty$, when $N_p = 5$ (solid line), $N_p = 15$ (dashed line), $N_p = 25$ (dotted line). The relaxation time is about 15 600 s.

respect to P_n in the absence of leakage, while the corresponding polarization time becomes $T_{pol} = 910$ s.

B. Competing *ISS* and electron spin lattice regime. The case of finite contact between electrons and nuclei, $T_{ISS} \approx T_{1e}$, is now considered.

In Fig. 6 and 7 build-up and relaxation curves, in the absence of leakage $(T_{1n} \rightarrow \infty)$ for $T_{1e} = 1$ s and T_{ISS} within the range 0.001 – 1 s, are analyzed. Two regimes can be identified. For a fast *ISS* process both T_{pol} and T_{relax} are not affected by the particular value of T_{ISS} , since the bottleneck process is represented by the electron spin-lattice relaxation. Conversely, when T_{ISS} grows the *ISS* process becomes the rate determining step and both polarization and relaxation



Fig. 5 Polarization build-up curves at 3.35 T and 1.2 K of a DNP system with $N_n/N_e = 1000$, $T_{1e} = 1$ s, $T_{ISS} = 0.001$ s for different nuclear intrinsic relaxation times: $T_{1n} \rightarrow \infty$ (solid line), $T_{1n} = 10000$ s (dashed line), $T_{1n} = 1000$ s (dot-dashed line), $T_{1n} = 100$ s (dotted line).



Fig. 6 Polarization build-up curves at 3.35 T and 1.2 K of a typical DNP system with $N_n/N_e = 1000$, $T_{1e} = 1$ s, $T_{1n} \rightarrow \infty$ for different values of contact time $T_{\rm ISS}$ (0.001 s (thick solid line), 0.01 s (thick dashed line), 0.1 s (thick dot-dashed line), 0.2 s (thin solid line), 0.5 s (thin dashed line), 1 s (thin dot-dashed line)).

times are enhanced. With $T_{\rm ISS} = 0.1$ s for instance, one obtains $T_{\rm pol} = 5300$ s and $T_{\rm relax} = 44\,900$ s. The ratio between the two characteristic times remains constant and close to 10. Finally no influence on the final polarization $P_{\rm n} = 0.825$ is observed on varying $T_{\rm ISS}$.

Surprisingly, when an almost negligible leakage term is introduced ($T_{1n} = 10\,000$ s), the final nuclear polarization becomes strongly dependent on the effectiveness of the contact between nuclear and electron systems (Fig. 8). For $T_{ISS} = 0.1$ s, the final nuclear polarization is $P_n = 0.529$, which corresponds to an important reduction of 36% that could be relevant for explaining some experimental observations. The polarization time is measured to be equal to 3200 s.

A comprehensive summary of the role played by the different parameters is presented in Fig. 9 and 10. Two distinct regimes



Fig. 7 Relaxation curves at 3.35 T and 1.2 K of a DNP system with $N_n/N_e = 1000$, $T_{1e} = 1$ s, $T_{1n} \rightarrow \infty$ for different values of contact time T_{ISS} (0.001 s (thick solid line), 0.01 s (thick dashed line), 0.1 s (thick dot-dashed line), 0.2 s (thin solid line), 0.5 s (thin dashed line), 1 s (thin dot-dashed line)).



Fig. 8 Polarization build-up curves at 3.35 T and 1.2 K of a DNP system with $N_n/N_e = 1000$ and $T_{1e} = 1$ s for different values of contact time T_{1SS} (0.001 s (thick lines) and 0.1 s (thin lines)) and nuclear intrinsic relaxation times T_{1n} ($\rightarrow \infty$ (solid lines), 10000 s (dashed lines) and 1000 s (dot-dashed lines)).

are clearly observable: highly and poorly effective contact between nuclear and electron systems. The former $(T_{\rm ISS} \rightarrow 0)$ is characterized by high levels of polarization and short build-up times, both dependent on $T_{\rm 1e}$, whilst the latter $(T_{\rm ISS} \rightarrow \infty)$ has low levels of polarization and long build-up times, both independent on $T_{\rm 1e}$. Clearly, in the first regime, the bottleneck role is played by $T_{\rm 1e}$, while $T_{\rm ISS}$ is sufficiently high not to affect the system dynamics. In the second regime their role is reversed.

V. Discussion and conclusion

To date, the main attempt to give a theoretical description of the DNP phenomenon at low temperature in the TM regime



Fig. 9 Final polarization *versus* contact time $T_{\rm ISS}$ for different values of the electron relaxation time T_{1e} (2 s (circles), 1 s (squares) and 100 ms (triangles)). Remaining parameters are set as follows: $N_n/N_e = 1000$, $T_{1n} = 10\,000$ s. Dashed lines represent polarization final values as estimated by the Borghini prediction in eqn (24).

has been the prediction proposed by Borghini^{3,22} for the steady state nuclear polarization that makes use of the standard parameterization:

$$P_{\rm n} = \tanh(\beta_{\rm n}\omega_{\rm n}) \quad \text{with } \beta_{\rm n} = \hbar/(2k_{\rm B}T_{\rm n}), \quad (22)$$

where T_n represents the temperature reached by the nuclear reservoir under microwave irradiation. To compute this temperature, the original derivation conjectures that the final electron polarization profile $(P_{e,i} = P_{e,i}(t \rightarrow \infty))$ takes the form:

$$P_{e,i} = -\tanh[\beta(\Delta_i - c)].$$
(23)

Actually, when the spectral diffusion is the fastest process, this form can be rigorously derived – without any *ad hoc* conjecture – by imposing the detailed balance condition as shown in Appendix A. In addition to eqn (23), a perfect thermal contact between the nuclear and the electron system and an infinite microwave power were assumed in the approach proposed by Borghini, so that $\beta = \beta_n$ and $c = \Delta_0$.

Under these assumptions, the celebrated relation (re-derived in Appendix C for the convenience of the reader):

$$\sum f_i (\Delta_i - \Delta_0) P_{e,i} + \Delta_0 P_0 - \omega_n \frac{N_n T_{1e}}{N_e T_{1n}} P_n = 0 \qquad (24)$$

gives a unique solution for T_n and thus for P_n . In the absence of leakage, the value of P_n predicted by eqn (24) at given temperature and field depends only on the ESR line shape and the irradiation frequency. For a Gaussian shape the maximal enhancement corresponds to $\omega_0 = \omega_e - \sigma$, where σ is the standard deviation of the electron frequency distribution.

The Borghini prediction is in good qualitative agreement with many experimental observations obtained in low temperature DNP experiments. However it does not provide an accurate quantitative description of experimental data and in particular of the maximum final nuclear polarization that turns always out to be overestimated.^{23,28–30} When a typical



Fig. 10 Polarization build-up times as a function of the contact time $T_{\rm ISS}$ for different values of the electron relaxation time $T_{\rm 1e}$ (2 s (circles), 1 s (squares) and 100 ms (triangles)). Remaining parameters are set as follows: $N_{\rm n}/N_{\rm e} = 1000$, $T_{\rm 1n} = 10000$ s.

trityl doped sample, with an EPR line width of 63 MHz, $T_{1e} = 1$ s, and $N_n/N_e = 1000$ is considered, eqn (24) predicts a maximum steady state polarization $P_n = 0.825$ in the absence of leakage. The same equilibrium value is obtained by our mathematical approach (Fig. 3 and 6) under the same physical constraints.

On the other hand the polarization levels achieved experimentally are in general much lower (see for instance ref. 23, 27 and 31). To justify this discrepancy, one can call on dissipation terms. In fact, if a finite T_{1n} is considered in eqn (24), a loss of P_n is obtained, that however, with T_{1n} generally being longer than 10^4 s,²³ can hardly exceed 10-15% (cfr. Fig. 5). A severe reduction of P_n is obtained instead by means of the equation set (18), introducing a finite contact between nuclei and electrons in the presence of a small leakage (see Fig. 8). It is worth noting that in the absence of leakage, *i.e.* when nuclei are isolated by the lattice, the final polarization is not affected by the efficiency of the contact between electron and nuclear systems.

Besides providing a more flexible scenario for P_n , the proposed mathematical framework allows the computation of polarization and relaxation times. Two regimes have been identified.

When the contact between nuclei and electrons is highly efficient, the bottleneck of the spin dynamics is T_{1e} and, by setting the parameters of the simulation according the experimental conditions used in ref. 23 one obtains $T_{\rm pol} \approx 10^3$ s and $T_{\rm relax} \approx 10^4$ s, as effectively measured in ref. 23. It is interesting to observe that, in this fast exchange limit, an estimation of the order of magnitude of $T_{\rm pol}$ and $T_{\rm relax}$ can be heuristically derived as follows. During polarization, the nuclear system transfers energy to electrons, which are cooled by the lattice. In a time T_{1e} , the lattice can absorb an energy proportional at most to $N_{\rm e}$, so that $T_{\rm pol} \approx N_{\rm n} T_{\rm 1e}/N_{\rm e} \approx$ 10^3 s, with $N_n/N_e = 1000$. During relaxation, the effective number of electrons that can absorb energy from the nuclear system, proportional to $(1 - P_{e,i})(1 + P_{e,i}) \approx (1 - P_0^2)$, collapses to ≈ 0.1 due to the high electron polarization $P_{e,i} \approx P_0 = 0.95$ at 3.35 T and 1.2 K, thus explaining the factor $T_{\rm relax}/T_{\rm pol} \approx 10$ observed both in simulations and experiments. On decreasing the effectiveness of the contact between nuclear and electron spins, a different regime is established,

where the *ISS* process becomes the rate determining step and both T_{pol} and T_{relax} become longer. This provides a possible explanation of why T_{pol} can range between very different values in samples with the same ratio N_e/N_n polarized under analogous conditions. By way of example, one can compare the [1-¹³C]-pyruvic acid samples doped with 15 mM of the trityl radical studied in ref. 23 and 27, where $N_e/N_n = 1000$ and $T_{pol} = 1200$ s with the [1-¹³C]-butyric acid sample mixed with 20% in volume of DMSO and doped with the 10 mM trityl radical analyzed in ref. 31, having a much longer $T_{pol} =$ 3400 s while the ratio N_e/N_n remains almost unchanged.

In conclusion, we propose a novel approach based on rate equations for studying the dependency of dynamic nuclear polarization in the low temperature thermal mixing regime $(T_{2e} \ll T_{1e})$ from the microscopic transitions involving electron and nuclear spins. This approach allows the recovery of the whole build-up curve and, in the limit of perfect contact between nuclei and electrons and infinite microwave power, leads to the same final nuclear polarization predicted by Borghini. In addition, by tuning the efficiency of the exchange interaction between nuclei and electrons, different values of P_n are reached, providing an interpretation key for those experimental observations of P_n which are not simply accounted for by leakage terms depending only on spin concentration and spin–lattice relaxation times.

The rate equation approach can be easily extended to more complex experimental systems. A second nuclear reservoir which also participates in TM could *e.g.* be included to interpret the dynamic experimental data measured in nitroxyl doped samples (where both ¹³C and ¹H Larmor frequencies do not exceed the ESR line width^{24,28,29}) or, similarly, in trityl doped samples containing ¹³C and ⁸⁹Y nuclei, both in contact with the electron reservoir.^{25,30} Finally the versatility of the approach proposed here would easily allow us to introduce new dissipative processes violating the precise assumptions and conservation principles the Borghini prediction is based on and that could possibly be useful to justify the many unexplained observations of low temperature DNP, *e.g.* the reduction of P_n on increasing N_e .^{27,32}

Appendix A

Detailed balance

After defining $P_{e,i}^+$ as the fraction of electrons up and $P_{e,i}^-$ as the fraction of electrons *down* belonging to the packet *i*, the detailed balance condition of the process depicted in Fig. 2, panel B, is:

$$P_{e,i-\delta}^{+}(P_{e,i}^{-})^{2} P_{e,i+\delta}^{+} = P_{e,i-\delta}^{-} (P_{e,i}^{+})^{2} P_{e,i+\delta}^{-}.$$
 (A1)

Then, by using the relation

$$P_{e,i} = \frac{2}{2}$$

 $P_{e,i}^{-} = \frac{(1 - P_{e,i})}{2}$

 $P^+ - \frac{(1+P_{e,i})}{2}$

one comes to an equation for the electron polarization

$$(1 + P_{e,i-\delta}) (1 - P_{e,i})^2 (1 + P_{e,i+\delta})$$

= $(1 - P_{e,i-\delta}) (1 + P_{e,i})^2 (1 - P_{e,i+\delta}),$ (A2)

that can be solved in the continuum limit where $\delta \omega \to 0$, $N_{\rm p} \to \infty$ and $P_{{\rm e},i} \to P_{{\rm e}}(\Delta_i)$. It is sufficient to consider the transitions between consecutive packets, so that $P_{{\rm e},i+\delta} = P_{{\rm e},i+1} \to P_{{\rm e}}(\Delta_i + \delta \omega)$, and write a second order expansion

$$P_{\rm e}(\Delta_i + \delta\omega) \approx P_{\rm e}(\Delta_i) + \delta\omega P'_{\rm e}(\Delta_i) + \frac{\delta\omega^2}{2} P''_{\rm e}(\Delta_i).$$
 (A3)

By combining relations (A2) and (A3), one gets the second order differential equation

$$2P_{\rm e}(\varDelta_i)P_{\rm e}'(\varDelta_i)^2 + P_{\rm e}''(\varDelta_i)(1 - P_{\rm e}(\varDelta_i)^2) = 0, \qquad (A4)$$

whose general solution

$$P_{\rm e}(\Delta_i) = -\tanh(\beta(\Delta_i - c)) \tag{A5}$$

was used in our treatment as a starting point for what we called the "STT" process. The same parametric function has been postulated in the derivation of the Borghini prediction for the steady state.

Appendix B

Microscopic derivation of rate equations

The term proportional to $1/T_{ISS}$ and describing the *ISS* process, used in the rate equations (eqn (18)), is here derived for the electron polarization first and then for nuclear polarization.

1. Electron polarization. Be $P_{e,i}^+$ the fraction of electrons up belonging to the packet *i*, P_n^+ the fraction of nuclei up, $P_{e,i}^-$ the fraction of electrons *down* belonging to the packet *i* and P_n^- the fraction of nuclei *down*. When the *ISS* event depicted in Fig. 11 occurs, the fraction of electrons up in the *i*-th packet is reduced by $1/(N_ef_i)$. The number of possible transitions is the product of:

• the number of the electrons up in the *i*-th packet: $N_{ef_i}P_{e,i}^+$, • the number of the electrons down in the $(i + \delta n_p)$ -th packet: $N_{ef_{i+\delta n_p}}P_{e,i+\delta n_p}^-$,

• the number of nuclei *down*: $N_{\rm n}P_{\rm n}^-$.

The rate of such a process is $1/(T_{\rm ISS}N_{\rm e}N_{\rm n})$, and the total reduction of $P_{\rm e,i}^+$ in the time interval dt is:

$$-\frac{\mathrm{d}t}{T_{\mathrm{ISS}}}f_{i+\delta n_{\mathrm{p}}}P_{\mathrm{e},i}^{+}P_{\mathrm{e},i+\delta n_{\mathrm{p}}}^{-}P_{\mathrm{n}}^{-}.$$
 (B1)

The total variation of $P_{e,i}^+$ induced by all possible *ISS* transitions, $\delta P_{e,i}^+$, is given by:

$$\delta P_{e,i}^{+} = \frac{dt}{T_{ISS}} \Big[f_{i+\delta n_{p}} (P_{e,i}^{-} P_{e,i+\delta n_{p}}^{+} P_{n}^{+} - P_{e,i}^{+} P_{e,i+\delta n_{p}}^{-} P_{n}^{-}) \\ + f_{i-\delta n_{p}} (P_{e,i}^{-} P_{e,i-\delta n_{p}}^{+} P_{n}^{-} - P_{e,i}^{+} P_{e,i-\delta n_{p}}^{-} P_{n}^{+}) \Big].$$



Fig. 11 Schematic representation of one possible ISS event.

Using the relations:

$$P_{e,i}^{+} = \frac{(1+P_{e,i})}{2}, \quad P_{e,i}^{-} = \frac{(1-P_{e,i})}{2}$$
$$P_{n}^{+} = \frac{(1+P_{n})}{2}, \quad P_{n}^{-} = \frac{(1-P_{n})}{2},$$

the total variation of $P_{e,i}$ induced by all possible *ISS* processes, $\delta P_{e,i} = 2\delta P_{e,i}^+$, can be written as follows:

$$\begin{split} \delta P_{\mathrm{e},i} &= \frac{\mathrm{d}t}{4T_{\mathrm{ISS}}} \left\{ f_{i+\delta n_{\mathrm{p}}} \left[(1-P_{\mathrm{e},i})(1+P_{\mathrm{e},i+\delta n_{\mathrm{p}}})(1+P_{\mathrm{n}}) \right. \\ &\left. - (1+P_{\mathrm{e},i})(1-P_{\mathrm{e},i+\delta n_{\mathrm{p}}})(1-P_{\mathrm{n}}) \right] \\ &\left. + f_{i-\delta n_{\mathrm{p}}} \left[(1-P_{\mathrm{e},i})(1+P_{\mathrm{e},i-\delta n_{\mathrm{p}}})(1-P_{\mathrm{n}}) \right. \\ &\left. - (1+P_{\mathrm{e},i})(1-P_{\mathrm{e},i+\delta n_{\mathrm{p}}})(1-P_{\mathrm{n}}) \right] \right\}. \end{split}$$
(B2)

The term proportional to $1/T_{ISS}$ in the first equation of set (18) can be now easily derived from eqn (B2) by means of simple algebraic calculations.

2. Nuclear polarization. When the event depicted in Fig. 11 occurs, the fraction of nuclei up is increased by a factor of $1/N_n$. The number of possible transitions is the product of:

• the number of the electrons up in the *i*-th packet: $N_e f_i P_{e,i}^+$,

• the number of the electrons down in the $(i + \delta n_p)$ -th packet: $N_e f_{i+\delta n_p} P_{e,i+\delta n_p}^-$,

• the number of nuclei down: $N_n P_n^-$.

The rate of such process is $1/(T_{ISS}N_eN_n)$, and the relevant increment of P_n^+ in the time interval dt is:

$$\frac{N_{\rm e} dt}{N_{\rm n} T_{\rm ISS}} f_i f_{i+\delta n_{\rm p}} P^+_{{\rm e},i} P^-_{{\rm e},i+\delta n_{\rm p}} P^-_{\rm n}.$$
 (B3)

Considering now all the possible processes, the total variation of P_n^+ induced by the *ISS* process, δP_n^+ , is given by:

$$\delta P_{\mathrm{n}}^{+} = \frac{N_{\mathrm{e}} dt}{N_{\mathrm{n}} T_{\mathrm{ISS}}} \sum_{i} f_{i} f_{i+\delta n_{\mathrm{p}}} [P_{\mathrm{e},i}^{+} P_{\mathrm{e},i+\delta n_{\mathrm{p}}}^{-} P_{\mathrm{n}}^{-} - P_{\mathrm{e},i}^{-} P_{\mathrm{e},i+\delta n_{\mathrm{p}}}^{+} P_{\mathrm{n}}^{+}].$$
(B4)

Following the line previously described for $P_{e,i}$, one immediately arrives at the equations for P_n reported in the main text (eqn (18)).

Appendix C

Borghini relation

To facilitate the reading of the manuscript, we report the derivation of the Borghini relation in eqn (24) according to the line proposed in ref. 3. The energy of the whole electron and nuclear system is conveniently split into two reservoirs: the Zeeman electron contribution

$$E_{Z_{e}}(t) = \frac{1}{2} N_{e} \hbar \omega_{e} \sum f_{i} P_{e,i}(t)$$
(C1)

and the non-Zeeman electron plus Zeeman nuclear term

$$E_{\text{NZ}-Z_{\text{n}}}(t) = \frac{1}{2} N_{\text{n}} \hbar \omega_{\text{n}} P_{\text{n}}(t) - \frac{1}{2} N_{\text{e}} \hbar \sum \Delta_{i} f_{i} P_{\text{e},i}(t). \quad (C2)$$

The time evolution of the two energy reservoirs is described by the following equations:

$$\frac{\mathrm{d}E_{Z_{\mathrm{c}}}(t)}{\mathrm{d}t} = \frac{1}{2}N_{\mathrm{e}}\hbar\omega_{\mathrm{e}}\sum f_{i}\frac{\mathrm{d}P_{\mathrm{e},i}(t)}{\mathrm{d}t}\frac{\mathrm{d}E_{\mathrm{NZ}-Z_{\mathrm{n}}}(t)}{\mathrm{d}t}$$

$$= \frac{1}{2}N_{\mathrm{n}}\hbar\omega_{\mathrm{n}}\frac{\mathrm{d}P_{\mathrm{n}}(t)}{\mathrm{d}t} - \frac{1}{2}N_{\mathrm{e}}\hbar\sum\Delta_{i}f_{i}\frac{\mathrm{d}P_{\mathrm{e},i}(t)}{\mathrm{d}t},$$
(C3)

and since the spectral diffusion and the *ISS* process conserve both $E_{Z_a}(t)$ and $E_{NZ-Z_a}(t)$, one obtains:

$$\frac{dE_{Z_{e}}(t)}{dt} = \frac{1}{2} N_{e} \hbar \omega_{e} \left[\sum f_{i} \frac{P_{0} - P_{e,i}(t)}{T_{1e}} - f_{0} \frac{P_{e,0}}{T_{1MW}} \right]$$
$$\frac{dE_{NZ-Z_{n}}(t)}{dt} = \frac{1}{2} N_{e} \hbar \left[\sum \Delta_{i} f_{i} \frac{P_{e,i}(t)}{T_{1e}} + f_{0} \Delta_{0} \frac{P_{e,0}}{T_{1MW}} \right]$$
$$-\frac{1}{2} N_{n} \hbar \omega_{n} \frac{P_{n}(t)}{T_{1n}}.$$
(C4)

It is important to observe that the evolution of the two energy reservoirs depends on the time progression of all $P_{e,i}(t)$ and $P_n(t)$, which is the full solution of the system of rate equations reported in eqn (18). As far as only the steady state solution is required, however, it is sufficient to impose the simultaneous vanishing of both right-hand sides of eqn (C4). Thus, by multiplying the first eqn (C4) by Δ_0/ω_e and adding it to the right-hand side of the second eqn (C4), one gets rid of the microwave transition probability $1/T_{1MW}$ and obtains the celebrated Borghini relation given in eqn (24).

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